

46

50

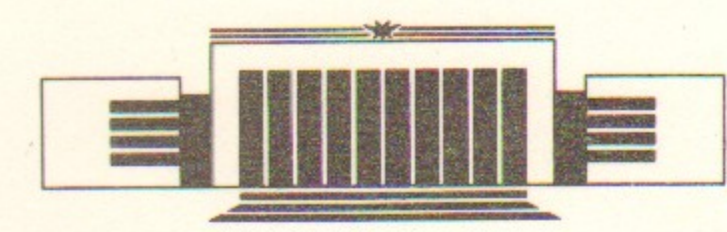


The State Scientific Center of Russia
The Budker Institute of Nuclear Physics
SB RAS

R.Zh. Shaisultanov

BACKREACTION IN SCALAR QED,
LANGEVIN EQUATION AND
DECOHERENCE FUNCTIONAL

Budker INP 95-82



НОВОСИБИРСК

Backreaction in scalar QED, Langevin equation and Decoherence functional

R.Zh. Shaisultanov

Budker Institute of Nuclear Physics
630090, Novosibirsk 90, Russia

Abstract

Using the Schwinger-Keldysh (closed time path or CTP) and Feynman-Vernon influence functional formalisms we obtain a Langevin equation for the description of the charged particle creation in electric field and of backreaction of charged quantum fields and their fluctuations on time evolution of this electric field. We obtain an expression for the influence functional in terms of Bogoliubov coefficients for the case of quantum electrodynamics with spin 0 charged particles. Then we derive a CTP effective action in semiclassical approximation and its cumulant expansion. An intimate connection between CTP effective action and decoherence functional will allow us to analyze how macroscopic electromagnetic fields are "measured" through interaction with charges and thereby rendered classical.

©The State Scientific Center of Russia
The Budker Institute of Nuclear Physics

1 Introduction

Nonequilibrium aspects of quantum field theory are beginning to receive considerable attention in recent years. Among them are dissipation and decoherence in quantum cosmology [1], structure formation in inflationary cosmology [2, 3, 4] and time evolution of a nonequilibrium chiral phase transition [5] to name only a few.

An interesting problem of similar nature is a problem of backreaction of particle production in a dynamical background field. Backreaction of quantum processes like particle creation in cosmological spacetimes [6] has been considered by many authors to understand how quantum effects affect the structure and dynamics of the universe near Planck time [7]. E. Calzetta and B.L. Hu [8] used Schwinger-Keldysh (or closed-time-path, CTP) functional formalism to derive a real and causal equation of motion (the semiclassical Einstein equation) describing the backreaction of particle production on classical effective geometry. In this equation one can identify a nonlocal kernel in the dissipative term whose integrated dissipative power has been shown to be equal to the energy density of the total number of particles created, thus establishing the dissipative nature of the backreaction process. Then in [9] Hu pointed out that a Langevin-type equation is what should be expected, and predicted that for quantum fields a colored noise source should appear in the driving term. Calzetta and Hu [10] show how noise and fluctuations can be attained with the CTP formalism together with dissipation and decoherence. Hu and Matacz [11] used cumulant expansion of the influence

functional of Feynman and Vernon, derived in terms of the Bogoliubov coefficients, to extract the noise associated with the matter field and to derive an Einstein-Langevin equation. In the Feynman-Vernon formalism [12, 13, 14] the effects of noise and dissipation can be extracted from the imaginary and real parts of the influence functional. Also it was shown [9, 15] that backreaction effect on the dynamics of space-time can be viewed as a manifestation of fluctuation-dissipation relation. Thus Hu and collaborators have extended the old framework of semiclassical gravity, based on Einstein equation with expectation value of energy-momentum tensor as a source, to that based on a Langevin equation which describes also the fluctuations of matter fields and spacetime (see recent review [16]).

In this article we will study the quantum non-equilibrium effects of pair creation in strong electric fields. Backreaction of pair creation on electric field was recently discussed by Cooper, Mottola et al [17]. They derived the semiclassical Maxwell equation, carry out its renormalization and numerically solve it for some initial conditions in 1+1 dimensions. Their numerical results clearly exhibits the decay of the electric field because of screening by the produced particles. We wish to make a step further and derive a Langevin equation, taking into account a noise from quantum matter fields. To this purpose we will use some mixture of Schwinger-Keldysh (CTP) and Feynman-Vernon influence functional formalisms. It is important to note here that phenomenological equations of motion with noise term can also be derived using decoherence functional formulation of quantum mechanics. This was done for some model quantum systems in [18, 19]

This paper is organized as follows: In Sec.2 we give a brief review of CTP functional formalism, mainly to introduce notations. All essential details can be found in [16, 20, 21, 22]. In Sec.3 we will obtain an expression for the influence functional in terms of Bogoliubov coefficients for the case of quantum electrodynamics with spin 0 charged particles. Then in Sec.4 we will obtain a CTP effective action in semiclassical approximation and its cumulant expansion. It is the main result of this paper. We will apply it for study of two interesting problems. First this result will allow us to analyze the backreaction of created charged particles on electric field and to derive a Langevin equation, which take into account a noise from quantum matter fields. Then in Sec.5 we will use an intimate connection between CTP effective action and decoherence functional [10, 16] to analyze how macroscopic electromagnetic fields are "measured" through interaction with charges and thereby rendered classical.

2 The Closed Time Path Functional Formalism in Quantum Field Theory

Usually in quantum field theory our interest is in obtaining the amplitudes of transition from in-states to the out-states. But in many cases, mainly in statistical physics, we are concerned with expectation values of physical quantities at finite time. To solve such initial value problems Schwinger has invented close time path (CTP) formalism.

Let us consider the expectation value of an arbitrary operator A :

$$\langle A \rangle (t) = \text{Tr } \rho(t) A \quad (1)$$

Here ρ is the density matrix that describes the (mixed) state of the system. The density matrix does not necessarily have to commute with the Hamiltonian, in which case it describes a non-equilibrium state. Using relation $\rho(t) = U(t,0) \rho(0) U^{-1}(t,0)$ where $U(t,0)$ is the evolution operator, inserting the identity operator $1 = U(t,T) U(T,t)$ we obtain

$$\langle A \rangle (t) = \text{Tr } \rho(0) U^{-1}(t,0) A U(t,0) = \text{Tr } \rho(0) U(0,T) U(T,t) A U(t,0) \quad (2)$$

Equation (2) can be pictured as describing the evolution of the system from 0 to t , inserting the operator A , evolving further to some large time T (in practice, $T \rightarrow \infty$), and then backwards from T to 0. The insertion of operator may be achieved by introducing external sources coupled to the particular operator. This suggests the definition of the CTP generating functional

$$Z[J_+, J_-] \equiv \exp iW[J_+, J_-] = \text{Tr } \rho(0) U(0, T, J_-) U(T, 0, J_+) \quad (3)$$

In the path integral representation we have

$$Z[J_+, J_-] = \int D\phi_1 D\phi_2 D\phi \langle \phi_1 | \rho | \phi_2 \rangle \int_{\phi_1}^{\phi} D\phi_+ \times \int_{\phi_2}^{\phi} D\phi_- \exp i \int_0^T dt \{L[\phi_+] - L[\phi_-] + J_+ \phi_+ - J_- \phi_-\} \quad (4)$$

The expectation values can be obtained as

$$\bar{\phi}_+ = \frac{\delta W}{\delta J_+}, \quad \bar{\phi}_- = \frac{\delta W}{\delta J_-} \quad (5)$$

Then the CTP effective action is

$$\Gamma_{CTP}[\bar{\phi}_+, \bar{\phi}_-] = W[J_+, J_-] - J_+ \bar{\phi}_+ + J_- \bar{\phi}_- \quad (6)$$

The equations of motion are

$$\frac{\delta \Gamma_{CTP}}{\delta \bar{\phi}_+} = -J_+, \quad \frac{\delta \Gamma_{CTP}}{\delta \bar{\phi}_-} = J_- \quad (7)$$

The physical situations correspond to solutions of the homogeneous equations at $\bar{\phi}_+ = \bar{\phi}_-$. Then equations are real and causal.

To apply this formalism to our situation we should substitute the ϕ field by the pair ψ and σ . We will be interested in expectation values of ψ only, so we do not couple the σ field to an external source. Also we assume that the initial density matrix factorizes $\rho = \rho_\psi \rho_\sigma$. Then we have

$$\begin{aligned} Z[J_+, J_-] &= \int D\psi_1 D\sigma_1 D\psi_2 D\sigma_2 \langle \psi_1 | \rho_\psi | \psi_2 \rangle \langle \sigma_1 | \rho_\sigma | \sigma_2 \rangle \\ &\int D\psi D\sigma \int_{\psi_1}^{\psi} D\psi_+ \int_{\sigma_1}^{\sigma} D\sigma_+ \int_{\psi_2}^{\psi} D\psi_- \int_{\sigma_2}^{\sigma} D\sigma_- \exp i \int_0^T dt \{ L_\psi[\psi_+] - \\ &- L_\psi[\psi_-] + J_+ \psi_+ - J_- \psi_- + L_\sigma[\sigma_+] - L_\sigma[\sigma_-] + L_{int}[\psi_+, \sigma_+] - \\ &- L_{int}[\psi_-, \sigma_-] \} = \int D\psi_1 D\psi_2 \langle \psi_1 | \rho_\psi | \psi_2 \rangle \int D\psi \int_{\psi_1}^{\psi} D\psi_+ \int_{\psi_2}^{\psi} D\psi_- \times \\ &\exp i \int_0^T dt \{ L_\psi[\psi_+] - L_\psi[\psi_-] + J_+ \psi_+ - J_- \psi_- \} \Phi[\psi_+, \psi_-] \end{aligned} \quad (8)$$

where $\Phi[\psi_+, \psi_-]$ is the so called influence functional

$$\begin{aligned} \Phi[\psi_+, \psi_-] &\equiv \exp i S_{IF}[\psi_+, \psi_-] = \int D\sigma_1 D\sigma_2 \langle \sigma_1 | \rho_\sigma | \sigma_2 \rangle \int_{\sigma_1}^{\sigma} D\sigma_+ \int_{\sigma_2}^{\sigma} D\sigma_- \\ &\exp i \int_0^T dt \{ L_\sigma[\sigma_+] - L_\sigma[\sigma_-] + L_{int}[\psi_+, \sigma_+] - L_{int}[\psi_-, \sigma_-] \} = \\ &\text{Tr} [U(T, 0; \psi_+) \rho_\sigma(0) U^{-1}(T, 0; \psi_-)] \end{aligned} \quad (9)$$

It is now easy to show, using (8) and (9), that in semiclassical approximation CTP effective action has the form

$$\Gamma_{CTP}[\psi_+, \psi_-] = S[\psi_+] - S[\psi_-] + S_{IF}[\psi_+, \psi_-] \quad (10)$$

From this relation one may derive the semiclassical equations of motion for the expectation values of the ψ field. It is worth noting that

$$\Gamma_{CTP}[\psi_+, \psi_-] = -\Gamma_{CTP}^*[\psi_-, \psi_+] \quad \text{and} \quad \Gamma_{CTP}[\psi, \psi] \equiv 0. \quad (11)$$

3 Influence functional for scalar QED

In this section we wish to find the influence functional in terms of Bogoliubov coefficients (as in [11]). Since we are dealing here with the case of scalar QED the pair of fields A and φ will play a similar role as ψ and σ in Sec.2 respectively. The influence functional have now the form

$$\Phi[A', A] = \text{Tr} [U(T, 0; A') \rho_\varphi(0) U^{-1}(T, 0; A)] \quad (12)$$

To obtain $U(T, 0; A)$ we will use the Heisenberg equation of motion

$$\ddot{\varphi} + \left(-i \vec{\nabla} - e \vec{A} \right)^2 \varphi + m^2 \varphi = 0 \quad (13)$$

We will choose $\vec{A} = (0, 0, A(t))$ and

$$\varphi(\vec{x}, t) = \sum_{\vec{p}} \frac{1}{\sqrt{2\omega_0(\vec{p})V}} \left\{ a_{\vec{p}}(t) + b_{-\vec{p}}^+(t) \right\} e^{i\vec{p}\vec{x}} \quad (14)$$

where $a_{\vec{p}}$ and $b_{-\vec{p}}^+$ are usual annihilation and creation operators for particles and antiparticles respectively. In what follows we will set volume $V = 1$. Here $\omega_0(\vec{p}) \equiv \omega(\vec{p}, 0)$ with

$$\omega^2(\vec{p}, t) = \vec{p}^2 + (p^3 - eA(t))^2 + m^2 \quad (15)$$

It is well known (see [23]) that the solution of (13) for creation and annihilation operators has the following form

$$a_{\vec{p}}(t) = \alpha_{\vec{p}}(t) a_{\vec{p}}(0) + \beta_{\vec{p}}^*(t) b_{-\vec{p}}^+(0); \quad b_{-\vec{p}}^+(t) = \beta_{\vec{p}}(t) a_{\vec{p}}(0) + \alpha_{\vec{p}}^*(t) b_{-\vec{p}}^+(0) \quad (16)$$

where the Bogoliubov coefficients $\alpha_{\vec{p}}(t)$ and $\beta_{\vec{p}}(t)$ are expressed via auxiliary function $\eta_{\vec{p}}(t)$ by following relations

$$\begin{aligned} \alpha_{\vec{p}}(t) &= \frac{1}{2\omega_0(\vec{p})} \left[i \frac{d\eta_{\vec{p}}(t)}{dt} + \omega_0(\vec{p}) \eta_{\vec{p}}(t) \right] \\ \beta_{\vec{p}}^*(t) &= \frac{1}{2\omega_0(\vec{p})} \left[i \frac{d\eta_{\vec{p}}^*(t)}{dt} + \omega_0(\vec{p}) \eta_{\vec{p}}^*(t) \right] \end{aligned} \quad (17)$$

with the $\eta_{\vec{p}}(t)$ satisfying to equation

$$\frac{d^2}{dt^2}\eta_{\vec{p}}(t) + \omega^2(\vec{p}, t)\eta_{\vec{p}}(t) = 0; \eta_{\vec{p}}(t) = e^{-i\omega_0(\vec{p})t} \text{ at } t \rightarrow 0 \quad (18)$$

Using last equations it is easy to show that $\alpha_{\vec{p}}(t)$ and $\beta_{\vec{p}}(t)$ obey to a system of ordinary first order differential equations

$$\dot{\alpha}_{\vec{p}}(t) = -ih_{\vec{p}}(t)\alpha_{\vec{p}}(t) - ig_{\vec{p}}(t)\beta_{\vec{p}}(t); \dot{\beta}_{\vec{p}}(t) = ig_{\vec{p}}(t)\alpha_{\vec{p}}(t) + ih_{\vec{p}}(t)\beta_{\vec{p}}(t) \quad (19)$$

with

$$h_{\vec{p}}(t) = \frac{1}{2\omega_0(\vec{p})}(\omega^2(\vec{p}, t) + \omega_0^2(\vec{p})) ; g_{\vec{p}}(t) = \frac{1}{2\omega_0(\vec{p})}(\omega^2(\vec{p}, t) - \omega_0^2(\vec{p})) \quad (20)$$

Equations (17-20) enable one to consider the particle creation in a homogeneous electric field $E(t)$ having an arbitrary time dependence. For example the famous Schwinger results [24], describing pair creation in a constant electric field, can be found using this equations [23]. Note that the number of created particles in \vec{p} th mode is given by

$$n_{\vec{p}} = |\beta_{\vec{p}}|^2 \quad (21)$$

Now we have enough formulae to find $U(t, 0)$ using relations

$$\begin{aligned} a_{\vec{p}}(t) &= U^\dagger a_{\vec{p}} U = \alpha_{\vec{p}} a_{\vec{p}} + \beta_{\vec{p}}^* b_{-\vec{p}}^+ \\ b_{-\vec{p}}^+(t) &= U^\dagger b_{-\vec{p}}^+ U = \beta_{\vec{p}} a_{\vec{p}} + \alpha_{\vec{p}}^* b_{-\vec{p}}^+ \end{aligned} \quad (22)$$

here we take for brevity $a_{\vec{p}} \equiv a_{\vec{p}}(0); b_{-\vec{p}}^+ \equiv b_{-\vec{p}}^+(0)$.

We will skip the details of calculations and express our result in the following form $U = \prod_{\vec{p}} U_{\vec{p}}$ where $U_{\vec{p}}$ is (we drop the mode label)

$$U = S(r, \phi) R(\theta) \quad (23)$$

where

$$S(r, \phi) = \exp [r (e^{2i\phi} a^\dagger b^\dagger - e^{-2i\phi} a b)]; R(\theta) = \exp [i\theta (a^\dagger a + b^\dagger b + 1)] \quad (24)$$

S and R are called two-mode squeeze and rotation operators respectively. The parameters r, ϕ, θ are determined from the equations

$$\alpha = e^{i\theta} \cosh r; \beta = e^{i\theta - 2i\phi} \sinh r \quad (25)$$

This expression for $U(t, 0)$ may be useful in many situations. We may, for example, rather easily describe time evolution of density matrix $\rho(t)$ from arbitrary initial $\rho(0)$, more interesting initial states are: vacuum state, thermal equilibrium and coherent states. Recently the time evolution of density matrix at finite temperature was considered in [25] using the functional Schrodinger representation. In this paper we will deal only with vacuum initial state. Applying (12) and (23) we find that the influence functional with vacuum as an initial state is given by

$$\Phi[A', A] = \prod_{\vec{p}} \frac{1}{\alpha_{\vec{p}}'^* \alpha_{\vec{p}} - \beta_{\vec{p}}'^* \beta_{\vec{p}}} \quad (26)$$

4 Semiclassical CTP effective action and Langevin equation

We may now, using (10) and (26), obtain for Γ_{CTP} :

$$\begin{aligned} \Gamma_{CTP}[A', A] &= S[A'] - S[A] + S_{IF}[A', A] \\ \text{where } S_{IF}[A', A] &= i \sum_{\vec{p}} \ln [\alpha_{\vec{p}}'^* \alpha_{\vec{p}} - \beta_{\vec{p}}'^* \beta_{\vec{p}}] \end{aligned} \quad (27)$$

It is useful to introduce new variables as

$$\Xi = \frac{1}{2}(A' + A); \Delta = A' - A \quad (28)$$

and define [11]

$$C_n(t_1, \dots, t_n; \Xi_{t_1, 0}, \dots, \Xi_{t_n, 0}) \equiv \frac{1}{i^{n-1}} \frac{\delta^n S_{IF}[A', A]}{\delta \Delta(t_1) \dots \delta \Delta(t_n)} \Big|_{\Delta=0} \quad (29)$$

The notation of $C_1(t_1; \Xi_{t_1, 0})$ means C_1 is a function of t_1 and also a functional of Ξ between endpoints t_1 and 0. By virtue of (11) the C_n 's are real quantities. As shown in [11] the C_n 's can be interpreted as cumulants of stochastic force. We will ignore all cumulants with $n > 2$, because the cumulants are of order e^n . This will mean that we are making a Gaussian approximation to the noise. The Γ_{CTP} is now

$$\begin{aligned} e^{i\Gamma_{CTP}[A', A]} &= e^{iS[A'] - iS[A] + i \int_0^\infty d\tau_1 \Delta(\tau_1) C_1(\tau_1; \Xi_{\tau_1, 0})} \\ &\times e^{-d\tau_2 \Delta(\tau_1) \Delta(\tau_2) C_2(\tau_1, \tau_2; \Xi_{\tau_1, 0}, \Xi_{\tau_2, 0})} \end{aligned} \quad (30)$$

The last term in (30) may be recognized as a characteristic functional of a stochastic gaussian process and we rewrite it in a form

$$\exp \left\{ -\frac{1}{2} \int_0^\infty d\tau_1 \int_0^\infty d\tau_2 \Delta(\tau_1) \Delta(\tau_2) C_2(\tau_1, \tau_2; \Xi_{\tau_1,0}, \Xi_{\tau_2,0}) \right\} = \int D\xi P[\xi, \Xi] \exp \left\{ i \int_0^\infty d\tau \Delta(\tau) \xi(\tau) \right\} \quad (31)$$

where

$$P[\xi, \Xi] = \exp \left\{ -\frac{1}{2} \int_0^\infty d\tau_1 \int_0^\infty d\tau_2 \xi(\tau_1) C_2^{-1}(\tau_1, \tau_2; \Xi_{\tau_1,0}, \Xi_{\tau_2,0}) \xi(\tau_2) \right\} \quad (32)$$

is the distribution of colour noise ξ . We can use now $\Gamma_{CTP}[A', A]$ to obtain the semiclassical Langevin equations of motion

$$\ddot{A}(t) = C_1(t; \Xi_{t,0}) + \xi(t) \quad (33)$$

In our case we have from (27) and (29)

$$C_1(t; \Xi_{t,0}) = e \int \frac{d^3\vec{p}}{(2\pi)^3 \omega_0(\vec{p})} (p^3 - eA(t)) |\alpha_{\vec{p}}(t) + \beta_{\vec{p}}(t)|^2 \quad \text{and} \\ C_2(\tau_1, \tau_2; \Xi_{\tau_1,0}, \Xi_{\tau_2,0}) = \langle \xi(\tau_1) \xi(\tau_2) \rangle = \frac{e^2}{2} \sum_{\vec{p}} \frac{(p^3 - eA(\tau_1))(p^3 - eA(\tau_2))}{\omega_0^2(\vec{p})} \left\{ \eta_{\vec{p}}^2(\tau_1) \eta_{\vec{p}}^{*2}(\tau_2) + \eta_{\vec{p}}^{*2}(\tau_1) \eta_{\vec{p}}^2(\tau_2) \right\} \quad (34)$$

The $C_1(t; \Xi_{t,0})$ is divergent and we must renormalize (33). As in [17] we will use adiabatic regularization method (or n-wave regularization of Zel'dovich) [6, 26]. The advantage of using adiabatic regularization lies in the fact that we can remove divergences before summing over modes. Potentially divergent terms does not appear and all integrals are finite. This is especially useful in numerical computations. Representing $\eta_{\vec{p}}(t)$ in the form

$$\eta_{\vec{p}}(t) = \sqrt{\frac{\omega_0(\vec{p})}{\Omega_{\vec{p}}(t)}} \exp \left\{ -i \int_0^t \Omega_{\vec{p}}(t') dt' \right\} \quad (35)$$

we see, that to perform renormalization, we need only in few first terms in adiabatic expansion of Ω :

$$\frac{1}{\Omega} = \frac{1}{\omega} + \left[\frac{\ddot{\omega}}{4\omega^2} - \frac{3\dot{\omega}^2}{8\omega^3} \right] + \dots \quad (36)$$

We rewrite now (33) in the form

$$\ddot{A}(t) = e \int \frac{d^3\vec{p}}{(2\pi)^3} (p^3 - eA(t)) \left[\frac{1}{\Omega_{\vec{p}}(t)} - \frac{1}{\Omega(t)} + \frac{1}{\Omega(t)} \right] + \xi(t) = e \int \frac{d^3\vec{p}}{(2\pi)^3} (p^3 - eA(t)) \left[\frac{1}{\Omega_{\vec{p}}(t)} - \frac{1}{\Omega(t)} \right] - \ddot{A}(t) \frac{e^2}{48\pi^2} \ln \frac{\Lambda^2}{m^2} + \xi(t) \quad (37)$$

where

$\frac{1}{\Omega(t)} \equiv \frac{1}{\omega} + \left[\frac{\ddot{\omega}}{4\omega^2} - \frac{3\dot{\omega}^2}{8\omega^3} \right]$ and Λ is a cutoff in the transverse momentum integration.

Remembering now that in scalar QED the usual charge renormalization factor Z_3 is equal to

$$Z_3 = 1 - \frac{e^2}{48\pi^2} \ln \frac{\Lambda^2}{m^2}; \quad \text{and that } e_r = Z_3^{\frac{1}{2}} e; \quad A_r = Z_3^{-\frac{1}{2}} A \quad (38)$$

we obtain

$$\ddot{A}_r(t) = e_r \Psi[e_r A_r] + \xi_r \\ \text{with } \Psi[e_r A_r] \equiv e_r \int \frac{d^3\vec{p}}{(2\pi)^3} (p^3 - e_r A_r(t)) \left[\frac{1}{\Omega_{\vec{p}}(t)} - \frac{1}{\Omega(t)} \right]; \quad \xi_r = Z_3^{\frac{1}{2}} \xi \quad (39)$$

In C_2 we must also change e to e_r . So we obtain the finite renormalized Langevin equation (39) that describe the process of pair production in a spatially homogeneous electric field and the backreaction from this pairs on time evolution of the electric field. The solution of the Langevin equation is beyond the scope of the present paper. We plan to consider this solution in future. Notice that without noise term Eq.(39) is equal to the semiclassical Maxwell equation, obtained in [17].

5 Decoherence in scalar QED

In this section we will show, using the results of previous sections, how the programme of decoherence [18, 27] can be applied in the context of quantum electrodynamics in some detail. We then analyze an example where macroscopic electromagnetic fields are "measured" through interaction with charges and thereby rendered classical. This example was discussed recently by Kiefer [28] using different point of view.

In the consistent or decoherent histories formulation of quantum mechanics [29, 30, 31] the complete description of a coupled ψ, σ system is given in terms of fine-grained histories $\psi(t), \sigma(t)$. Let us take as a coarse-graining procedure of summing over the σ field. In other words the σ field play in our

case the role of environment. Then the interference effects between coarse-grained histories are measured by the decoherence functional $D[\psi, \psi']$. It was shown in [10, 16, 32] that the decoherence functional, which is the fundamental object of the decoherent histories formulation, is connected with CTP effective action by following relation

$$D[\psi, \psi'] = e^{i\Gamma_{CTP}[\psi, \psi']} \quad (40)$$

The coarse-grained history $\psi(t)$ can be described classically if and only if the decoherence functional is approximately diagonal, that is, $D[\psi, \psi'] \simeq 0$ whenever $\psi \neq \psi'$. Now after this very brief discussion of the some aspects of decoherence, we will proceed to discuss an example where macroscopic field strengths decohere through their interaction with charges. We wish to consider, as an example, a macroscopic superposition of two electric fields, one pointing upwards, and the other pointing downwards. In the case of scalar QED we have

$$D[A, A'] = e^{i\Gamma_{CTP}[A', A]} \sim e^{-\frac{1}{2} \int_0^\infty d\tau_1 \int_0^\infty d\tau_2 \Delta(\tau_1) \Delta(\tau_2) C_2(\tau_1, \tau_2; \Xi_{\tau_1, 0}, \Xi_{\tau_2, 0})} \quad (41)$$

Here we have omitted unessential phase factor. Now after an easy calculation we obtain

$$D[A, A'] \sim \exp \left\{ -\frac{V e^2 E^2}{256\pi m} \right\} \quad (42)$$

This result resembles the result obtained by Kiefer. Note that the interaction with the charge states leads to an exponential suppression factor of the corresponding interference terms for the field; in the infrared limit of $V \rightarrow \infty$ one finds exact decoherence. In realistic cases, however, a finite coherence width remains, so one can in principle subject these results to experimental confirmation. For an electric field of $E \approx 10^7$ Volts per centimeter, for example, one finds that interference effects are observable on length scales $L \leq 10^{-4}$ centimeters [28].

Thus the programme of decoherence [30] may successfully be applied in the context of quantum field theory using the concepts and methods of nonequilibrium statistical field theory.

Acknowledgments

I would like to thank V.N.Baier for his interest in this work and valuable comments on the text of the manuscript.

References

- [1] L.Hu, *Physica* A158,399 (1989); E.Calzetta and B.L.Hu, *Phys.Rev.* D35,495 (1987); D37,2838 (1988)
- [2] E.W.Kolb and M.S.Turner, *The early universe* (Addison-Wesley, Reading, MA, 1990)
- [3] A.Linde, *Particle physics and inflationary cosmology* (Harwood, New York, 1990)
- [4] A.A.Starobinsky, in *Field theory, quantum gravity and strings*, ed. H.J. de Vega and N.Sanchez (Springer, Berlin 1986)
- [5] F.Cooper, Y.Kluger, E.Mottola and J.P.Paz, *Phys.Rev.* D51, 2377(1995); D.Boyanovsky, H.J.de Vega, R.Holman, hep-th/9412052
- [6] L.Parker, *Phys.Rev.* 183, 1057 (1969); D3, 346 (1971); Ya.B.Zel'dovich, *Pis'ma Zh. Eksp. Teor. Fiz.* 12, 443(1970) [*JETP Lett.* 12, 307(1970)]; Ya.B.Zel'dovich and A.A.Starobinsky, *Zh. Eksp.Teor. Fiz.* 61, 2161(1971) [*Sov.Phys.JETP* 34, 1159(1972)]
- [7] Ya.B.Zel'dovich and A.A.Starobinsky, *Zh. Eksp.Teor. Fiz.* 61, 2161(1971) [*Sov.Phys.JETP* 34, 1159(1972)]; B.L.Hu and L.Parker, *Phys.Rev.* D17,933(1978); J.B.Hartle and B.L.Hu, *Phys.Rev.* D20, 1757(1979); D20, 1772(1979)
- [8] E.Calzetta and B.L.Hu, *Phys.Rev.* D35,495 (1987); D37,2838 (1988); D40, 656 (1989)
- [9] B.L.Hu, *Physica* A158,399 (1989)
- [10] E.Calzetta and B.L.Hu, *Phys.Rev.* D49, 6636(1994)
- [11] B.L.Hu and A.Matacz, "Backreaction in Semiclassical Cosmology: the Einstein-Langevin Equation", gr-qc/9403043, *Phys.Rev.* D51,(1995)
- [12] R.Feynman and F.Vernon, *Ann.Phys. (NY)* 24, 118 (1963); R.Feynman and A.Hibbs, *Quantum Mechanics and Path Integrals*, (McGraw-Hill, New York, 1965)
- [13] A.O.Caldeira and A.J.Leggett, *Physica* A121,587 (1983); *Ann.Phys. (NY)* 149,374 (1983)
- [14] H.Grabert, P.Schramm and G.L.Ingold, *Phys.Rep.* 168,115 (1988)

- [15] B.L.Hu and S.Sinha, "Fluctuation-Dissipation Relation in Cosmology", Univ.Maryland preprint pp93-164 (1993)
- [16] B.L.Hu, "Quantum Statistical Field Theory in Gravitation and Cosmology" in Proc. Third International Workshop on Thermal Field Theories and Applications, eds. R. Kobes and G. Kunstatter (World Scientific, Singapore, 1994)
- [17] F.Cooper and E.Mottola, Phys. Rev. D40,456(1989); F.Cooper, J.M.Eisenberg, Y.Kluger, E.Mottola and B.Svetitsky, Phys.Rev. D45,4659(1992); D48, 190(1993)
- [18] M.Gell-Mann and J.B.Hartle, Phys.Rev. D47,3345(1993)
- [19] T.A.Brun, Phys.Rev. D47,3383(1993)
- [20] J.Schwinger, J.Math.Phys. 2, 407(1961); L.V.Keldysh, Zh. Eksp. Teor. Fiz. 47, 1515 (1964) [Sov.Phys.JETP 20, 1018 (1965)]; G.Zhou, Z.Su, B.Hao and L.Yu, Phys.Rep. 118, 1 (1985);D.Boyanovsky, H.J.de Vega, R.Holman,hep-th/9412052
- [21] B.S. DeWitt, in Quantum Concepts in Space and Time, ed. R.Penrose and C.J.Isham (Clarendon Press, Oxford, 1986)
- [22] R.D.Jordan, Phys.Rev. D33, 44 (1986)
- [23] V.S.Popov, Zh.Eksp.Teor.Fiz. 62, 1248 (1972); M.S.Marinov and V.S.Popov, Forsch.Phys. 25, 373 (1977)
- [24] J.Schwinger, Phys.Rev.82, 664 (1951);Phys.Rev.93, 615 (1954)
- [25] J.Hallin and P.Liljenberg, "Fermionic and bosonic pair creation in an external electric field at finite temperature using the functional Schrodinger representation" preprint Goteborg ITP 94-40
- [26] N.Birrell and P.C.W.Davies, Quantum Fields in Curved Spaces (Cambridge University Press, Cambridge, 1982)
- [27] W.H.Zurek, Phys.Rev.D 24, 1516 (1981); Phys.Rev.D 26, 1862 (1982); in Frontiers of Nonequilibrium Statistical Physics, ed. G.T.Moore and M.O.Scully (Plenum, N.Y., 1986); Physics Today 44, 36 (1991); E. Joos and H.D.Zeh, Z.Phys. B59, 223 (1985); A.O.Caldeira and A.J.Leggett, Phys.Rev.A 31, 1059 (1985); W.G.Unruh and W.H.Zurek, Phys.Rev. D40, 1071 (1989); B.L.Hu,

- J.P.Paz and Y.Zhang, Phys.Rev.D45, 2843 (1992); Phys.Rev.D47, 1576 (1993); W.H.Zurek, J.P.Paz and S.Habib, Phys.Rev.Lett. 47, 1187 (1993); J.P.Paz, S.Habib and W.H.Zurek, Phys.Rev.D47, 488 (1993); J.P.Paz and W.H.Zurek, Phys.Rev.D48, 2728 (1993); J.Twamley, Phys.Rev. D48, 5730 (1993)
- [28] C. Kiefer, Phys. Rev. D 46, 1658 (1992);C. Kiefer,"Irreversibility in quantum field theory" preprint Freiburg THEP-95/2, quant-ph/9501004
- [29] R. B. Griffiths, J. Stat. Phys. 36, 219 (1984); R. Omnés, J. Stat. Phys. 53, 893, 933, 957 (1988); Ann. Phys. (NY) 201, 354 (1990); M. Gell-Mann and J. B. Hartle, in Complexity, Entropy and the Physics of Information, ed. by W. H. Zurek (Addison-Wesley, Reading, 1990). H. F. Dowker and J. J. Halliwell, Phys. Rev. D46, 1580 (1992). E. Calzetta and B. L. Hu, "Decoherence of Correlation Histories" in Directions in General Relativity, Vol II: Brill Festschrift, eds B. L. Hu and T. A. Jacobson (Cambridge University Press, Cambridge, 1993)
- [30] JB.Hartle, "Spacetime Quantum Mechanics and the Quantum Mechanics of Spacetime", Les Houches lectures (1992)
- [31] R.Omnès, Rev. Mod. Phys. 64, 339 (1992), W. H. Zurek, Prog. Theor. Phys. 89, 281 (1993)
- [32] J.P.Paz and S.Sinha,Phys.Rev.D45,2823(1992)